Wave physics

1D wave equation, travelling waves, sinusoidal waves and complex representation, standing waves, eigenmodes, electromagnetic waves in vacuum

Feynman Vol. II Chapter 20

Reminder from last lecture

Maxwell's equations

$$\operatorname{div} \vec{E} = \frac{\rho}{\epsilon_0}$$

Maxwell-Gauss equation

$$\operatorname{div} \vec{B} = 0$$

Absence of magnetic monopoles

$$\operatorname{curl} \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$

Maxwell-Faraday equation

$$\operatorname{curl} \vec{B} = \mu_0 \vec{j} + \mu_0 \epsilon_0 \frac{\partial \vec{E}}{\partial t}$$

Maxwell-Ampère equation

$$c = \frac{1}{\sqrt{\epsilon_0 \mu_0}} = 3.0 \times 10^8 \,\mathrm{m \, s^{-1}}$$

 $c=rac{1}{\sqrt{\epsilon_0\mu_0}}=3.0 imes10^8~\mathrm{m~s^{-1}}$ Light is no longer something else. It's the propagation of an electromagnetic disturbance according to the laws of electromagnetism.



Maxwell's equations describe electricity, magnetism and light in one single unified theory.

Examples	of waves
1. Examples of waves	
What is a wave in physics?	
A wave is a physical disturbance propagating in space or oscillating. Propagation is associated to energy transport, but no transport of matter constituents.	
Examples?	Physical Disturbance?
 Vibrating string (or 2D membrane) 	String vertical displacement
► Sound in air	Pressure, air density, longitudinal displacement
Earthquake (seismic waves)	Displacement
Water waves	Water displacement, surface elevation
► Gravitational waves	Spacetime metric defining distance
► Plasma waves	Charge density, electric field

• etc.

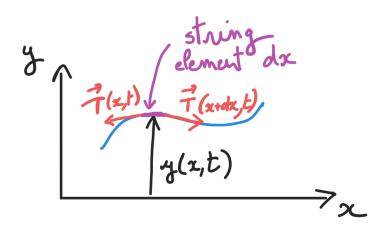
1. Examples of waves

Vibrating string

Mass per unit length: $\mu = dM/dx$

String tension: T = const

Small angles: $\tan \alpha = \frac{\partial y}{\partial x} \simeq \alpha$



Applying Newton's second law to a string element dx, we get (see tutorial #9):

$$\frac{\partial^2 y}{\partial x^2} - \frac{\mu}{T} \frac{\partial^2 y}{\partial t^2} = 0$$

y(x,t) is the string vertical displacement

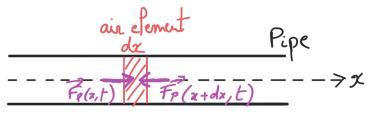
d'Alembert wave equation

1. Examples of waves

Sound in air, in one dimension

Small perturbations of pressure P and volume mass density μ :

$$P = P_0 + P_1$$
 with $P_1 \ll P_0$ $\mu_1 \ll \mu_0$ $\mu_1 \ll \mu_0$



Thermodynamics - adiabatic process:

$$PV^{\gamma} = \text{const} \implies P_1 = \kappa \mu_1 \quad \text{with} \quad \kappa = \frac{\partial P}{\partial \mu} = \gamma \frac{P_0}{\mu_0} \qquad \qquad \frac{\partial \mu_1}{\partial t} + \mu_0 \frac{\partial v}{\partial x} = 0 \quad \text{eq. (2)}$$

Mass conservation

Newton's second law for air element dx: $\mu_0 \frac{\partial v}{\partial t} = -\frac{\partial P_1}{\partial x}$ eq. (3)

Combining equations (1-3), we get:

$$\frac{\partial^2 P_1}{\partial x^2} - \frac{1}{\kappa} \frac{\partial^2 P_1}{\partial t^2} = 0$$

 $P_1(x,t)$ is the air pressure perturbation

d'Alembert wave equation

2. Solutions of 1D d'Alembert wave equation

D'Alembert wave equation in one dimension:

$$\frac{\partial^2 u}{\partial x^2} - \frac{1}{v^2} \frac{\partial^2 u}{\partial t^2} = 0$$

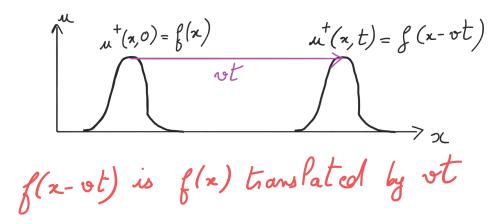
Linear equation: principle of superposition applies.

The general form of the solution of d'Alembert wave equation is:

$$u(x,t) = f(x - vt) + g(x + vt)$$

With f and g being arbitrary functions. Can be demonstrated by performing a change of variables to $\xi = x - vt$ and $\eta = x + vt$.

Travelling wave: $u^+(x,t) = f(x-vt)$



The wave solution u^+ is a signal travelling without any deformation at velocity v towards the right (in the direction of increasing x).



The constant v in d'Alembert wave equation is the speed of the wave.

Vibrating string Sound
$$v = \sqrt{\frac{T}{\mu}} \qquad v = \sqrt{\kappa} = \sqrt{\frac{\partial P}{\partial \mu}}$$

2. Solutions of 1D d'Alembert wave equation

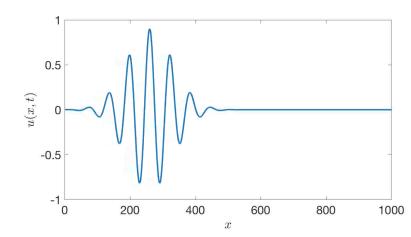
Second form for the travelling wave: $u^-(x,t) = g(x+vt)$ translation by -vt

The wave solution u^- is a signal travelling without any deformation at velocity v towards the left (in the direction of decreasing x).

Note: the general solution is not necessarily a travelling wave, but can always be written of the sum of wave u^+ travelling towards the right and of a wave u^- travelling towards the left.

Example of right-travelling wave

$$u^+(x,t) = f(x - vt)$$



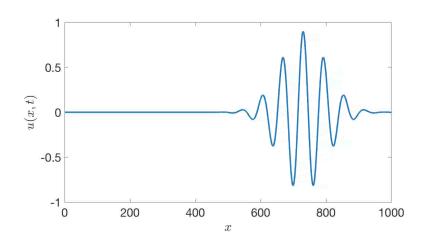
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Example of lefttravelling wave

$$u^{-}(x,t) = g(x+vt)$$



2. Solutions of 1D d'Alembert wave equation

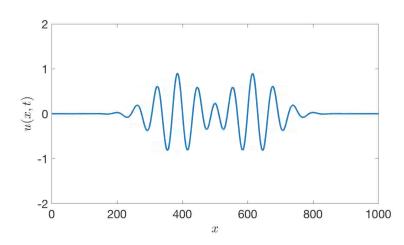
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Example of general solution

$$u(x,t) = f(x - vt) + g(x + vt)$$



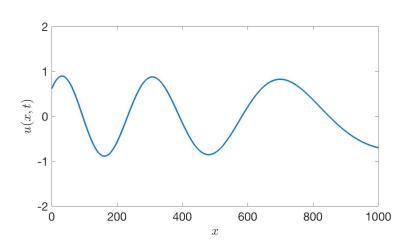
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More complex case

$$u^+(x,t) = f(x - vt)$$



2. Solutions of 1D d'Alembert wave equation

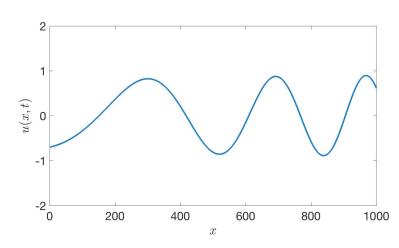
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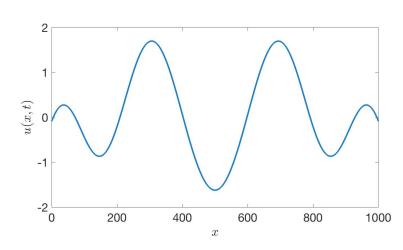
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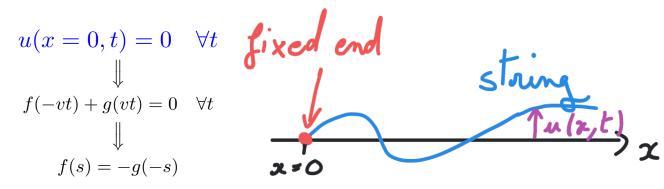
More complex case

$$u(x,t) = f(x - vt) + g(x + vt)$$



2. Solutions of 1D d'Alembert wave equation

What happens if there is a fixed end? —— There is a reflection.



Let's assume we have an incident wave coming from the right (x>0) and travelling to the left:

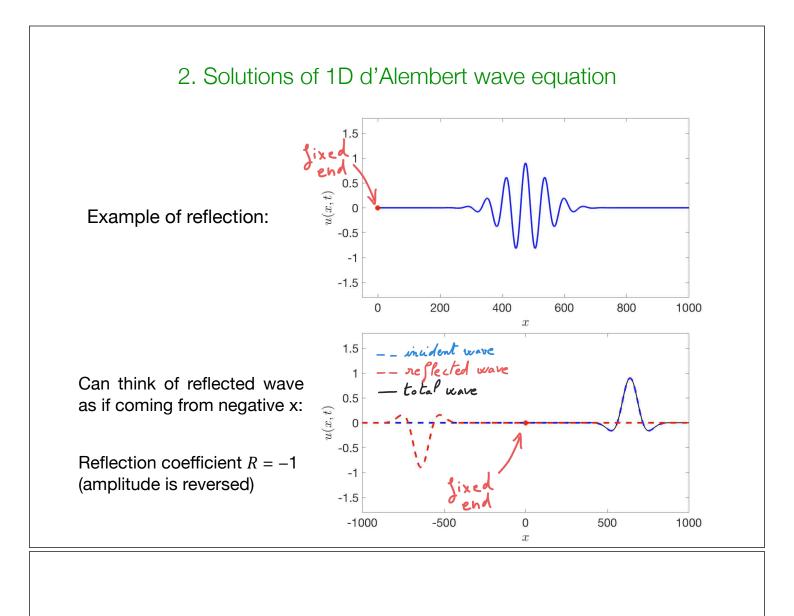
$$u_{\rm inc}(x,t) = g(x+ct)$$

The solution with fixed end at x=0 reads:

$$u(x,t) = f(x-vt) + g(x+vt) = -g(-(x-vt)) + g(x+vt)$$

reflected wave

incident wave



Sinusoidal waves

3. Sinusoidal waves

There is a preferred class of solutions called sinusoidal travelling waves:

$$u(x,t) = f(x-vt) = A\cos(kx-\omega t + \phi)$$
 right travelling

$$u(x,t) = g(x+vt) = A\cos(kx+\omega t + \phi)$$
 left travelling

with
$$\omega=kv$$

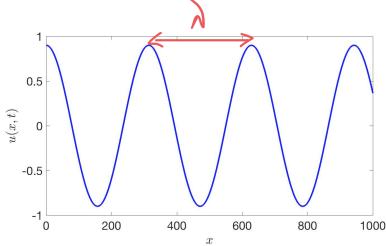
dispersion relation (relation between ω and k) for d'Alembert wave equation

$$\omega$$
 (angular) frequency $k=\omega/v$ wavenumber

$$T=rac{2\pi}{\omega}$$
 time period $\lambda=rac{2\pi}{k}=vT$ wavelength

3. Sinusoidal waves

right travelling sinusoidal wave $u(x,t) = A \, \cos(kx - \omega t)$



Complex representation (greatly simplifies calculation with sinusoidal functions):

$$u(x,t) = A \cos(kx - \omega t + \phi) \longrightarrow \underline{u}(x,t) = \underline{A} e^{i(kx - \omega t)}$$
 with $\underline{A} = A e^{i\phi}$

The real signal is recovered by taking the real part: $u(x,t) = \operatorname{Re}\left[u(x,t)\right]$

3. Sinusoidal waves

Why using sinusoidal waves?

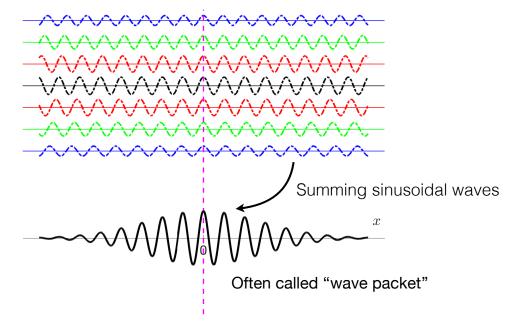
- ► Strictly speaking, a wave with infinite extent (to $x \to +\infty$ and $x \to -\infty$) and infinite energy has no physical reality.
- But it's a good approximation for many physical situations in which the signal is sinusoidal with a large extent (both in space and time).
- · Even more important: any real wave can be written as a sum (superposition) of sinusoidal waves

Sinusoidal waves form a base of the solution space. Translated in mathematics, it's the Fourier transform (note: discrete sum = Fourier series for periodic functions):

$$u(x,t) = \int_{-\infty}^{\infty} \underline{A}(k) \, e^{i(kx - \omega t)} \, dk \qquad + \underset{\text{with kx + wt continuous sum (superposition)}}{+}$$

3. Sinusoidal waves

Example of superposition of sinusoidal waves



Standing waves and eigenmodes

4. Standing waves and eigenmodes

Superposition of two counter-propagating sinusoidal waves of frequency ω of same amplitude:

$$u(x,t) = u_0 \cos(kx - \omega t) + u_0 \cos(kx + \omega t)$$
right travelling

$$\Longrightarrow u(x,t) = 2u_0 \cos(kx) \cos(\omega t)$$
 standing wave

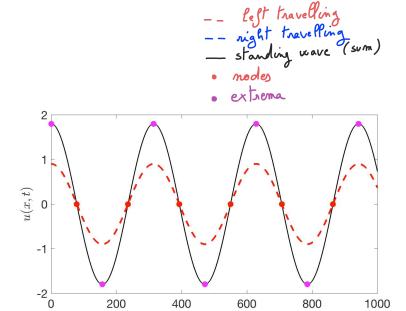
(does not travel, but oscillates)

Extrema (minimum and maximum): $kx_n=n\pi \implies x_n=n\frac{\lambda}{2}$

Nodes (u=0):
$$kx_n=\frac{\pi}{2}+n\pi$$
 \Longrightarrow $x_n=\frac{\lambda}{4}+n\frac{\lambda}{2}$

Trigonometric identity: $\cos{(A\pm B)} = \cos{A}\cos{B} \mp \sin{A}\sin{B}$

4. Standing waves and eigenmodes



standing wave $u(x,t) = 2u_0 \cos(kx) \cos(\omega t)$

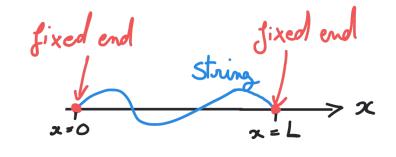
4. Standing waves and eigenmodes

What happens if there are two fixed end?

We look for standing wave solutions, who have fixed nodes separated by $\lambda/2$.

$$\implies L = N \frac{\lambda}{2}, \quad N \in \mathbb{N}^*$$

$$\implies k_N = \frac{2\pi}{\lambda} = \frac{N\pi}{L}$$



Standing waves are solutions only for wavenumbers that are integer multiple of π/L . They are called eigenmodes and read:

$$u_N(x,t) = \left(A\cos(k_N x) + B\sin(k_N x)\right) \left(C\cos(\omega_N t) + D\sin(\omega_N t)\right)$$
 with $\omega_N = k_N v$

$$u_N(0,t) = 0 \implies A = 0 \implies u_N(x,t) = \sin(k_N x)(C\cos(\omega_N t) + D\sin(\omega_N t))$$

eigenmodes

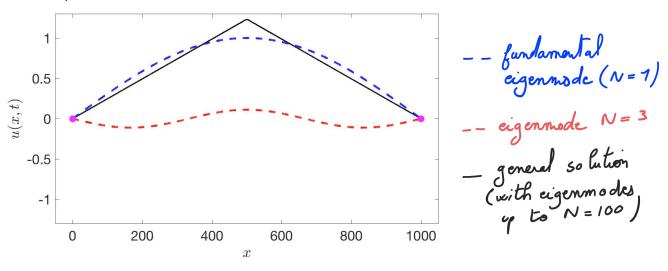
4. Standing waves and eigenmodes

What happens if there are two fixed end?

General solution as a discrete sum (superposition) over the eigenmodes:

$$u(x,t) = \sum_{N=1}^{+\infty} \sin(k_N x) \left[C_N \cos(\omega_N t) + D_N \sin(\omega_N t) \right]$$

Example:



1D electromagnetic wave in vacuum

5. 1D electromagnetic waves in vacuum

Electromagnetic fields in vacuum: $\rho = 0$ $\vec{j} = 0$

We assume the fields only depends on x (1D case):

$$\frac{\partial}{\partial y} = \partial_y = 0; \quad \frac{\partial}{\partial z} = \partial_z = 0 \quad \Longrightarrow \vec{\nabla} = \begin{bmatrix} \partial_x \\ 0 \\ 0 \end{bmatrix}$$

Under these assumptions, Maxwell's equations:

$$\frac{\partial E_x}{\partial x} = 0 \qquad \begin{vmatrix} \partial_x \\ 0 \\ 0 \end{vmatrix} \times \begin{vmatrix} E_x \\ E_y \\ E_z \end{vmatrix} = \begin{vmatrix} 0 \\ -\partial_x E_z \\ \partial_t B_y \\ \partial_t B_z \end{vmatrix}$$

Maxwell-Gauss equation

Maxwell-Faraday equation

$$\frac{\partial B_x}{\partial x} = 0 \qquad \begin{vmatrix} \partial_x \\ 0 \\ B_y \end{vmatrix} = \begin{vmatrix} 0 \\ -\partial_x B_z \\ \partial_x B_y \end{vmatrix} = \begin{vmatrix} \partial_t E_x \\ \partial_t E_y \\ \partial_t E_z \end{vmatrix}$$

Absence of magnetic monopoles

Maxwell-Ampère equation

5. 1D electromagnetic waves in vacuum

Equations for E_y and B_z :

$$\frac{\partial E_y}{\partial x} = -\frac{\partial B_z}{\partial t} \qquad \text{and} \qquad -\frac{\partial B_z}{\partial x} = \mu_0 \epsilon_0 \frac{\partial E_y}{\partial t}$$

$$\implies \frac{\partial^2 E_y}{\partial x^2} - \mu_0 \epsilon_0 \frac{\partial^2 E_y}{\partial t^2} = 0 \qquad \frac{\partial^2 B_z}{\partial x^2} - \mu_0 \epsilon_0 \frac{\partial^2 B_z}{\partial t^2} = 0$$

d'Alembert wave equation for
$$E_y$$
 and B_z with speed $c=\frac{1}{\sqrt{\mu_0\epsilon_0}}$

Maxwell's equations have wave solutions in vacuum travelling at c

Summary

1D d'Alembert wave equation:
$$\frac{\partial^2 u}{\partial x^2} - \frac{1}{v^2} \frac{\partial^2 u}{\partial t^2} = 0$$
 linear equation: principle of superposition applies

General solution:
$$u(x,t) = f(x-vt) + g(x+vt)$$

$$u^{+}(x,t) = f(x-vt)$$
 $u^{-}(x,t) = g(x+vt)$

right-travelling wave

left-travelling wave

Preferred class of solutions:
$$u(x,t) = A\cos(kx \pm \omega t + \phi)$$
 with $\omega = kv$ sinusoidal waves dispersion relation

Complex representation:
$$\underline{u}(x,t) = \underline{A} e^{i(kx\pm\omega t)}$$
 and $u(x,t) = \operatorname{Re}\left[\underline{u}(x,t)\right]$

Standing wave: 2 counter-propagating sinusoidal waves of same frequency and amplitude

One fixed end: reflection Two fixed ends: eigenmodes
$$k_N = \frac{N\pi}{L}$$
 $\omega_N = k_N v$

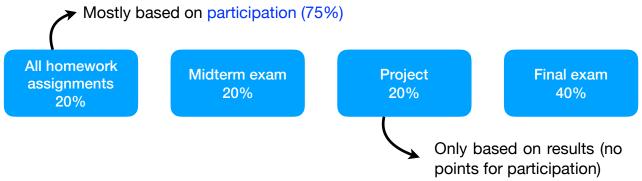
1D d'Alembert wave equation for *E* and *B* fields in vacuum:

$$\frac{\partial^2 E_y}{\partial x^2} - \mu_0 \epsilon_0 \frac{\partial^2 E_y}{\partial t^2} = 0 \qquad \frac{\partial^2 B_z}{\partial x^2} - \mu_0 \epsilon_0 \frac{\partial^2 B_z}{\partial t^2} = 0$$

Maxwell's equations have EM wave solutions in vacuum travelling at c

Reminder on grading system

Your final grade for PHY104 will be composed of:



Practical considerations for the project:

- You'll have 2 months to work on it, project due on 14 June 2018
- You can work in groups of up to 4 students, only one project to submit per group
- Topics explored: introduction to special relativity incl. time dilatation (10 points), $E=mc^2$ (5 bonus points), acceleration to near speed of light (10 points). Project grades above 20/20 (up to 25/20) will be accounted when computing the final grade.
- The project includes numerical work to solve a relativistic equation of motion.